

# Status of MERIS Calibration for 4<sup>th</sup> reprocessing

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- Instrument overview
- In-flight calibration hardware
- Overall calibration strategy
  - \* 3<sup>rd</sup> reprocessing radiometric validation results

#### Calibration processing chain

- \* Radiometric model
- Processing chain
- Why a temporal model?
- ✤ Gain modelling

#### Some results from whole mission reanalysis

- Reference diffuser stability
- Impact on gain drift determination

#### Conclusion



## Instrument overview

- Push-broom imaging spectrometer, 5 fan-arranged cameras
- CCD sensing with [390,1040] nm spectral range
- Up to 45 programmable µchannels relaxed into 15 channels





# In-flight Calibration hardware

- Calibration wheel with 5 positions:
- Shutter: dark offset (calibration zero), before every diffuser acquisition
- Radiometric diffuser: calibration gains (every 2 weeks)
- Reference radiometric diffuser: ageing of nominal diffuser (every 3 months)
- Spectral diffuser: spectral calibration at 2 wavelengths (every 3 months @520nm, every 6 months @ 408nm)
- All calibrations near orbital South pole, where Sun aligns with baffle





# **Overall Calibration Strategy**



DC = Dark Calibration RC = Radiometric Calibration WC = Wavelength Calibration

Orbital South Pole

Self standing absolute calibration for the EO processing chain

Validated by vicarious methods



#### **3**<sup>rd</sup> **RP Radiometry Validation results**

(from B. Fougnie, CNES, MERIS QWG, June 2013)

General comparison of all results from various methods

- ✦ Good consistency between all results for all methods
- Very good accordance for
  - » 412-443-490-510 within 1% between the 3 methods
  - » 753-778-865-885 within 0.5% between the 2 methods



#### Radiometric model



$$X_{b,k,m,t} = NL_{b,m} \begin{bmatrix} A_{b,k,m}^{0} \cdot (L_{b,k,m,t} + SL_{b,k,m,t} (L_{*,*,*,*})) + \\ Sm_{b,k,m,t} (L_{*,k,m,*} + SL_{*,k,m,*} (L_{*,*,*,*})) + \\ g_{C}(T_{t}^{CCD}) \cdot C_{b,k,m}^{0} \end{bmatrix} + \varepsilon$$

Where:

- b = band, k / m = pixel / camera, t = time, (\* = whole/partial domain)
- X<sub>b,k,m,t</sub> is the OLCI raw sample
- NL<sub>b,m</sub> is a non-linear function
- T<sup>CCD</sup>(t) is the temperature of the CCDs
- $g_C(T^{CCD})$  is a dimensionless temperature correction function
- A<sup>0</sup><sub>b,k,m</sub> the "absolute radiometric gain" in counts/radiance unit
- L<sub>b,k,m,t</sub> the spectral radiance distribution in front of OLCI
- $Sm_{b,k,m,t}$  the smear signal, due to continuous sensing of light by OLCI
- C<sup>0</sup><sub>b,k,m</sub> the calibrated dark signal (possibly including an on-board compensation)
- SL<sub>b,k,m,t</sub> a linear operator representing the stray light contribution to the signal
- ε is a random process representative of the noise and measurement errors.



### Radiometric model applies to Calibration measurements

$$\begin{split} X_{b,k,m,t}^{Cal} &= NL_{b,m} \begin{bmatrix} A_{b,k,m}^{0} \cdot \left( L_{b,k,m,t}^{Cal} + SL_{b,k,m,t} (L_{*,*,*,*}^{Cal}) \right) + \\ Sm_{b,k,m,t} (L_{*,k,m,*}^{Cal} + SL_{*,k,m,*} (L_{*,*,*,*}^{Cal})) + \\ g_{C} (T_{t}^{CCD}) \cdot C_{b,k,m}^{0} \end{bmatrix} + \mathcal{E} \\ A_{b,k,m}^{0} &= \left\langle \frac{NL_{b,m}^{-1} \left( X_{b,k,m,t}^{Cal} \right) - Sm_{b,k,m,t} (L_{*,k,m,*}^{Cal} + SL_{b,k,m,t}^{Cal}) - g_{C} (T_{t}^{CCD}) \cdot C_{b,k,m}^{0} \\ \left( L_{b,k,m,t}^{Cal} + SL_{b,k,m,t}^{Cal} \right) \right) \right\rangle \\ \end{split}$$

- X<sup>Cal</sup> from Sun diffuser measurements
- C<sup>0</sup> from dedicated measurements (with shutter)
- Sm from dedicated band (virtual, lit only during CCD frame transfer)
- L<sup>cal</sup> from characterised/modelled diffuser BRDF + in-flight geometry + E<sub>0</sub> at MERIS bands & pixels (from Spectral characterisation/model)
- SL from L + characterised/modelled convolution kernels
- gC from characterisation
- NL<sup>-1</sup> from characterisation

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## **Calibration chain summary**





# Calibration chain summary (short term)



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# Why using a Gain model

#### Allows smooth long-term trends correction

# BRDF model does not fully capture dependency with Sun azimuth (±0.5%)





On-ground: BRDF model residuals for 2 extreme SAA In-flight: Relative gain variations: extreme SAA gains over central one

#### Selecting "best" SAA ensures better accuracy.

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- Model is basically G(t) = G(t<sub>0</sub>)·f(t-t<sub>0</sub>)
- Correcting Diffuser 1 (and 2!) for ageing allows to measure consistent gains from both diffusers
- Gain long-term drift (Instrument Degradation) captured from D1 (more data, good spread in SAA domain)
- Validated on D2
- Gain at reference time can be derived from D1 or D2
  - D1 pros: minimize speckle
  - D2 pros: minimize discrepancies at camera interfaces

**Stability with time and view angle verified over Antarctica** 



# Some results, 1

- Ageing correction is essential for accurate gain drift modelling
- Assumption: ageing proportional to cumulated exposure (D2 exposed ~10 times less)
- Ageing measured by evolution of D1/D2 ratio, shows up to ~2% variation (in the blue)



#### ➔ D2 ageing no longer negligible



# Some results, 2

Neglecting D2 ageing has a small absolute impact (<0.2%) but degrades overall consistency: D1 and D2 do not "see" the same instrument degradation anymore: example of camera 2.





Some results, 3

#### Still work in progress, sorry!

# Here you should see an illustration of the impact of ageing revision 4<sup>th</sup> RP model, in particular over the 3<sup>rd</sup> RP extrapolation period.



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## Conclusion

Well characterized on-board diffusers are extremely accurate calibration "sources".

As for any space borne item ageing is a concern and shall be closely monitored.

#### A reference diffuser has proven to be a reliable monitoring device

- unfrequently exposed diffusers seem to degrade fairly linearly
- time sampling is important to minimize geometry effects
- But accurate ageing rate determination requires time





# Thank you for your attention