



SNPP VIIRS Reflective Solar Bands On-Orbit Calibration: Improvements, Updates and SDR Reprocessing

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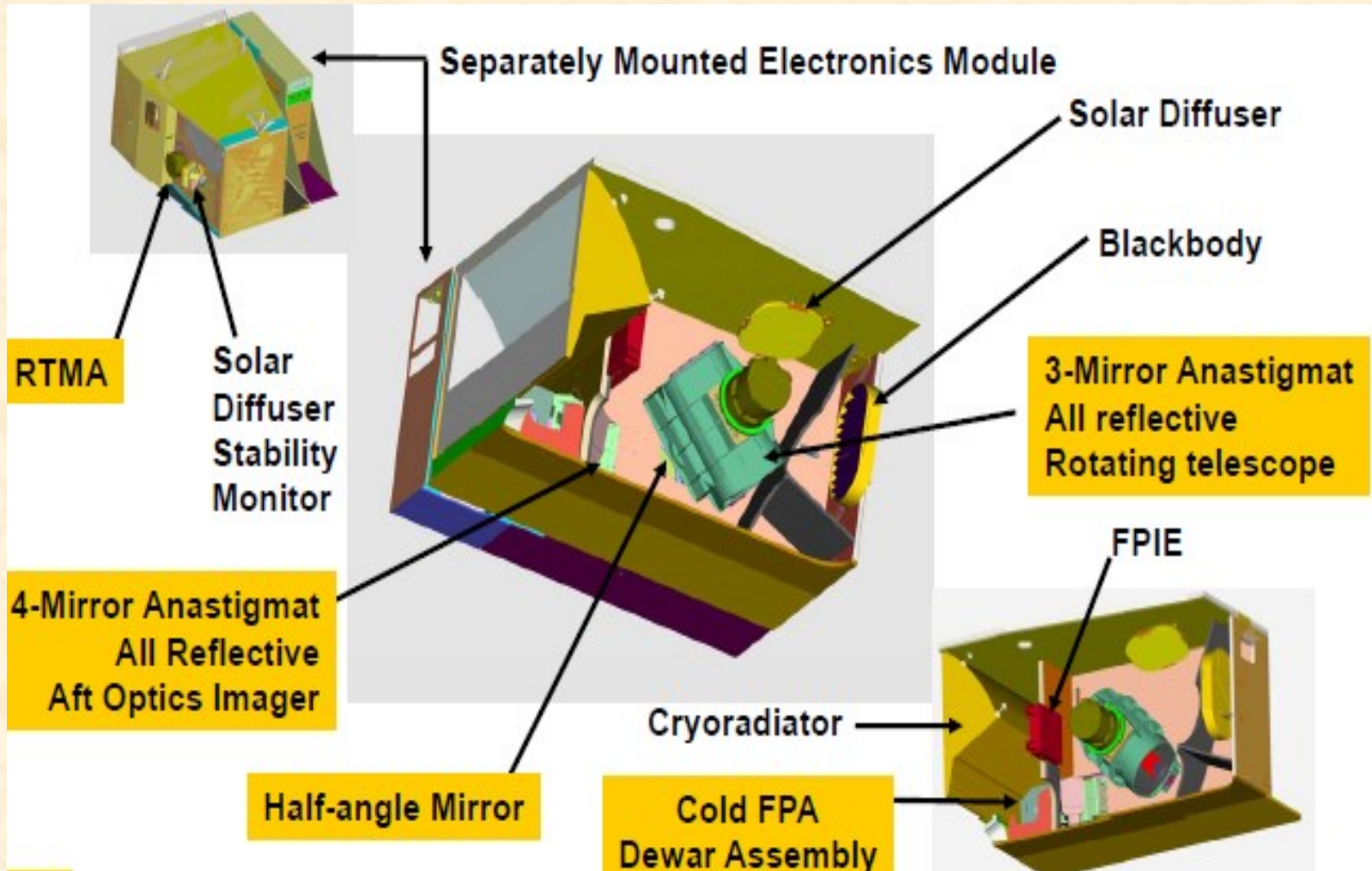


Outline



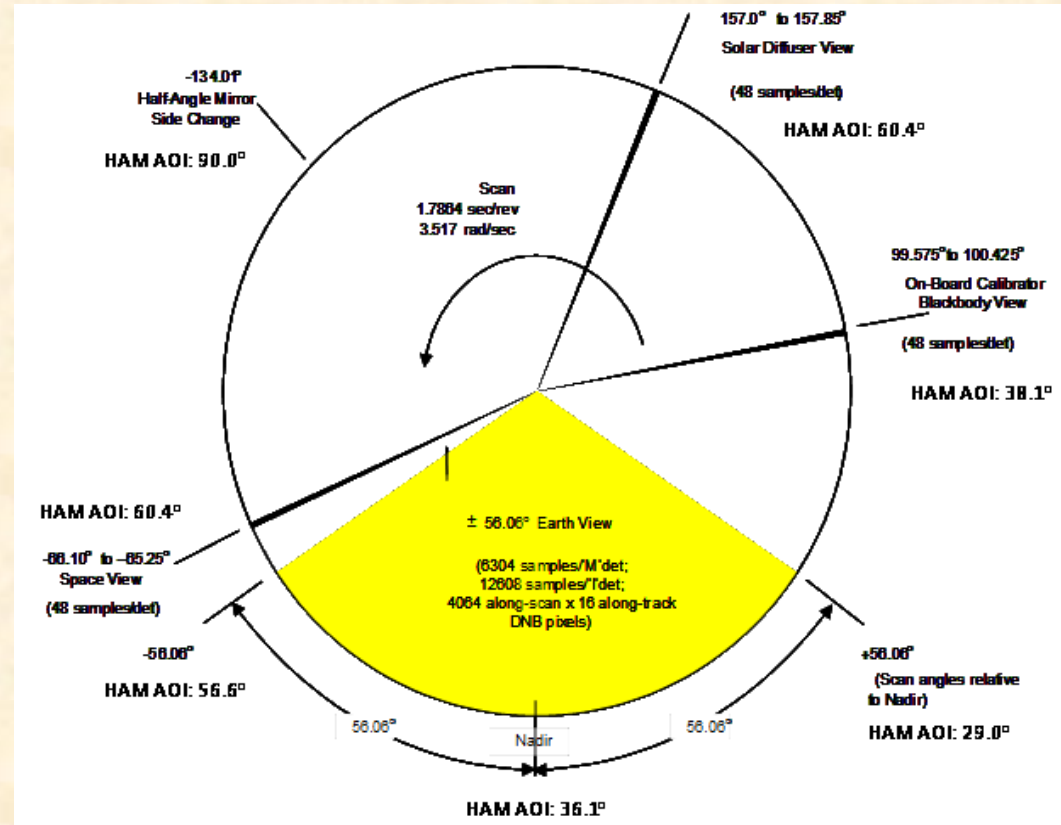
- Introduction
- Solar Diffusor Stability Monitor (SDSM) Calibration (H-factor)
 - Algorithms, data analysis, and performance
- Solar Diffusor (SD) Calibration (SD F-factor)
 - Algorithms, data analysis, and performance
- Lunar Calibration (Lunar F-factor)
 - Algorithms, data analysis, and performance
- Hybrid Approach (Hybrid F-factor)
 - Algorithms and hybrid calibration coefficients
- Inter-sensor and In-situ Comparison
- Ocean Color Products Performance
- Summary

VIIRS Background



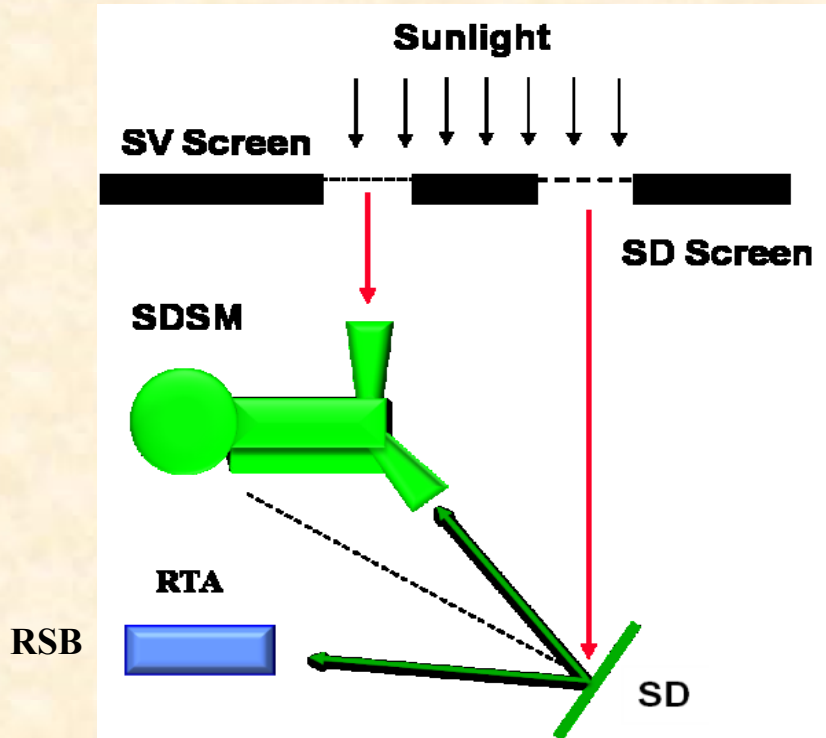
RSB On-Orbit Calibration

- 22 spectral bands - 410 nm to 12.013 μm spectral range
- 14 Reflective Solar Bands (RSB) : 3 image bands, I1-I3, and 11 moderate bands, M1-M11
- The VIIRS RSB are calibrated on orbit by SD/SDSM calibration
- Monthly lunar observation through its space view (SV) since launch.
- For VIIRS, the angle of incidence (AOI) of the SV is exactly the same as that of the SD. Lunar observations should provide identical on-orbit gain change for VIIRS RSB as SD/SDSM calibration.



VIIRS RSB uncertainty specification is 2%, but ocean color EDRs (using M1-M7, NIR; also M8, M10, and M11, NIR-SWIR; recently I1) need to achieve ~0.2%. This has been achieved.

SD/SDSM Calibration Overview

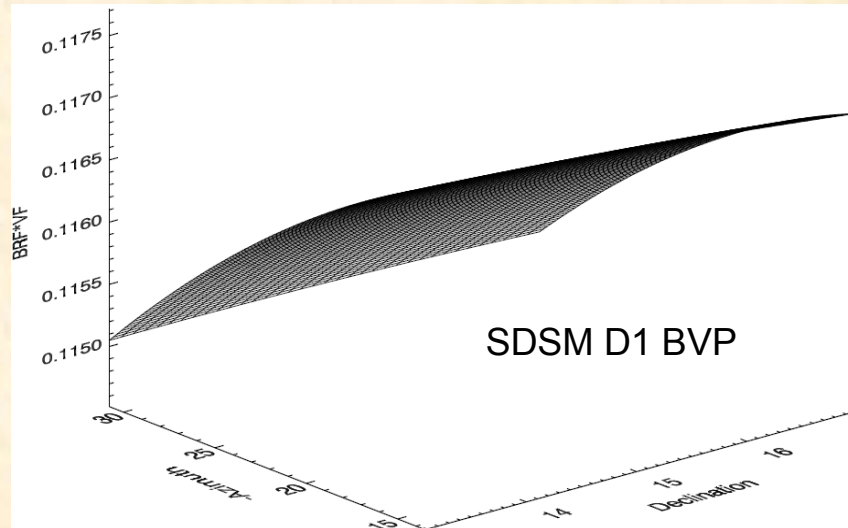
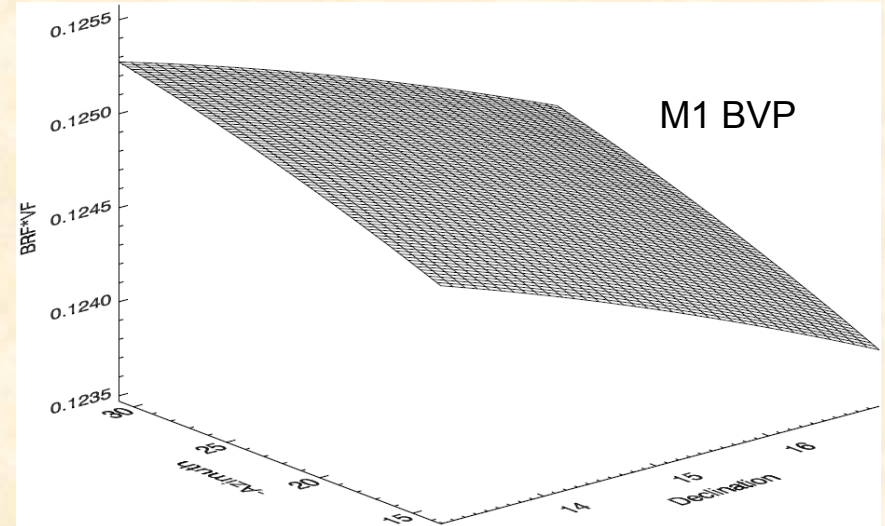
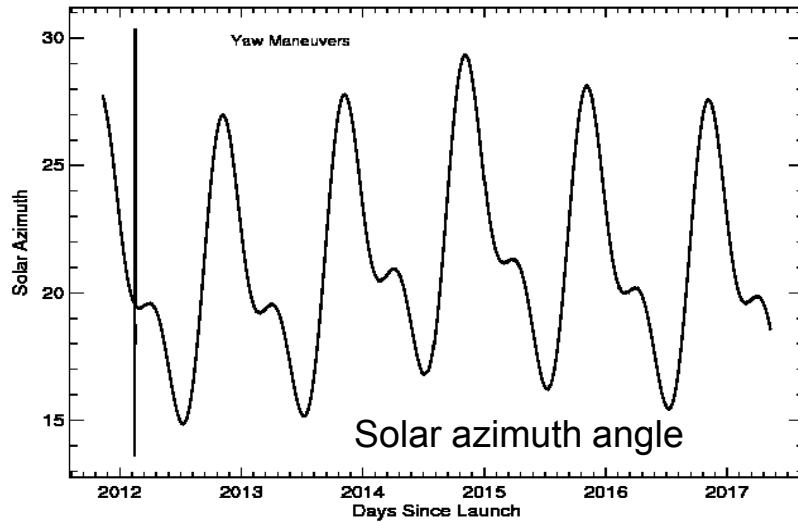


Key assumption: SD degrades uniformly with respect to both incident and outgoing directions

- *SD and SDSM sun view screens:*
 - *Prevent RSB and SDSM saturation*
 - *Vignetting functions (VFs)*
 - *VFs measured prelaunch and validated by yaw measurements*
 - *SD bidirectional reflectance factors (BRFs)*
- *BRFs measured prelaunch and validated by yaw measurements*
 - *SD on-orbit degradation is tracked by the SDSM measurements at 8 wavelength from 412 nm to 935 nm*
- *SDSM measures H-factors*
- *F-factors, or RSB calibration coefficients, are the final calibration product*

SD BRF and SDS VF

BRF-VF Product (BVP)



- Yaw carefully planned, cover all solar angle range for SD/SDSM calibration.
- Carefully derive BRFs and VFs from the yaw measurements is the crucial 1st step.
- **Need to do it right just one time from yaw data.**

J. Sun and M. Wang, "On-orbit characterization of the VIIRS solar diffuser and solar diffuser screen," Appl. Opt., 54, 236 -252 (2015).

SDSM Calibration Algorithm

- SDSM is a ratio radiometer, which views SD, Sun, and an internal dark scene successively in three-scan cycles.
- SD BRF for SDSM view direction

$$BRF_{SD,SDSM}(\lambda, t) = \rho_{SD,SDSM}(\lambda)H(\lambda, t)$$

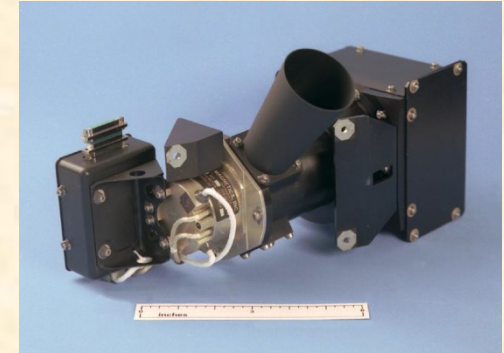
- $\rho_{SD,SDSM}(\lambda)$: Prelaunch BRF for SDSM view direction
 - $H(\lambda)$ is solar diffuser degradation since launch
- SD degradation, H factors, for SDSM view direction at the wavelength of the SDSM detector D

$$H(\lambda_D) = \left\langle \frac{dc_{SD,D}}{\rho_{SD,SDSM}(\lambda_D)\tau_{SDS} \cos(\theta_{SD})} \right\rangle_{Scan} \bigg/ \left\langle \frac{dc_{SV,D}}{\tau_{SVS}} \right\rangle_{Scan}$$

- **Improvements**

- **Robust and accurate VFs and BRFs from yaw measurements**
 - **Ratio of the averages**
 - **Sweet spots selection**

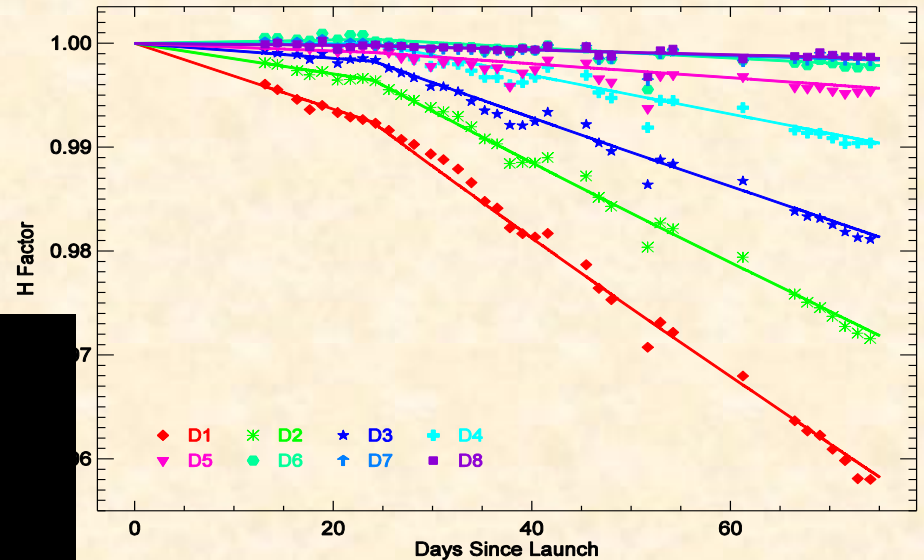
J. Sun and M. Wang, “Visible infrared image radiometer suite solar diffuser calibration and its challenges using solar diffuser stability monitor,” Appl. Opt., 53, 8571-8584 (2014).



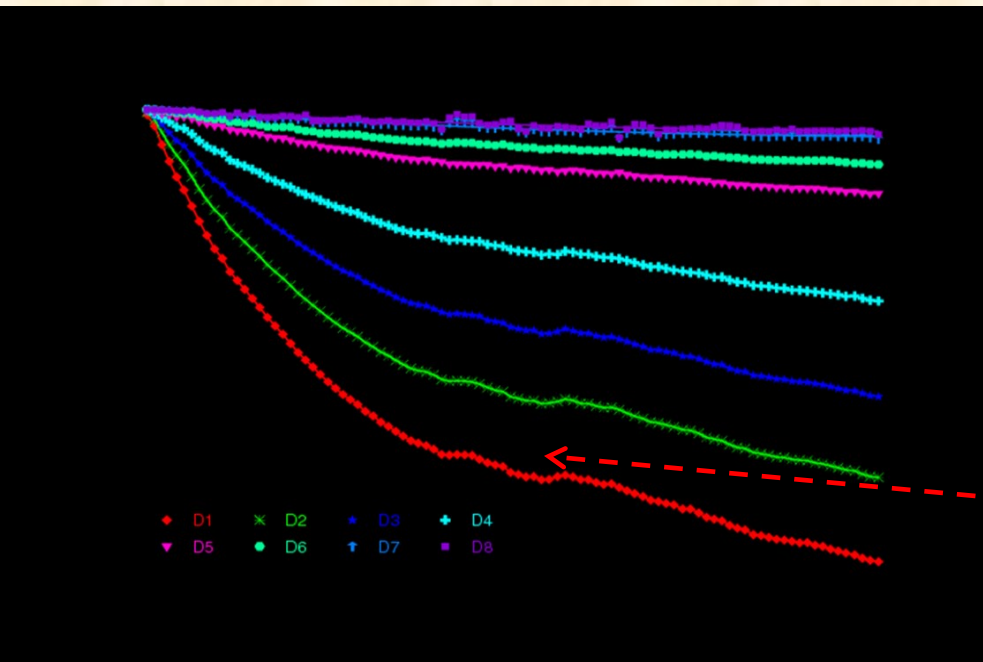
SDSM operations: Every orbit first few months, then once per day for about two years, and once per two days since May, 2014.

- *Very stable*
- *SDSM is very accurate!*
- *Results are all actual measurements*
- *No averaging over orbits, no smoothing, NO FANCY TRICKS*

SD Degradation – First 70 days



- *First 25 days behaved differently.*
- *But H-factor is in different direction from RSB view direction – KEY ISSUE*
- *Unexpected but real degradation features (Nov 2014)*



J. Sun and M. Wang, “Visible infrared image radiometer suite solar diffuser calibration and its challenges using solar diffuser stability monitor,” Appl. Opt., 53, 8571-8584 (2014).

- SD is made of Spectralon®, near Lambertian property
- Solar radinace reflected by the SD

$$L_{SD}(\lambda) = I_{Sun}(\lambda) \cdot \tau_{SDS} \cdot \cos(\theta_{SD}) \cdot \rho_{SD,RTA}(\lambda) \cdot h(\lambda) / d_{VS}^2$$

- $\rho_{RSD,RTA}(\lambda)$: Prelaunch BRF for RTA view direction
- $h(\lambda)$: SD degradation for SDSM view direction is used as the SD degradation for the RTA direction

- RSB calibration coefficients, F factors

$$F(B, D, M, G) = \frac{RVS_{B,SD} \cdot \int RSR_B(\lambda) \cdot L_{SD}(\lambda) \cdot d\lambda}{\sum_i c_i(B, D, M, G) \cdot dn^i \cdot \int RSR_B(\lambda) \cdot d\lambda}$$

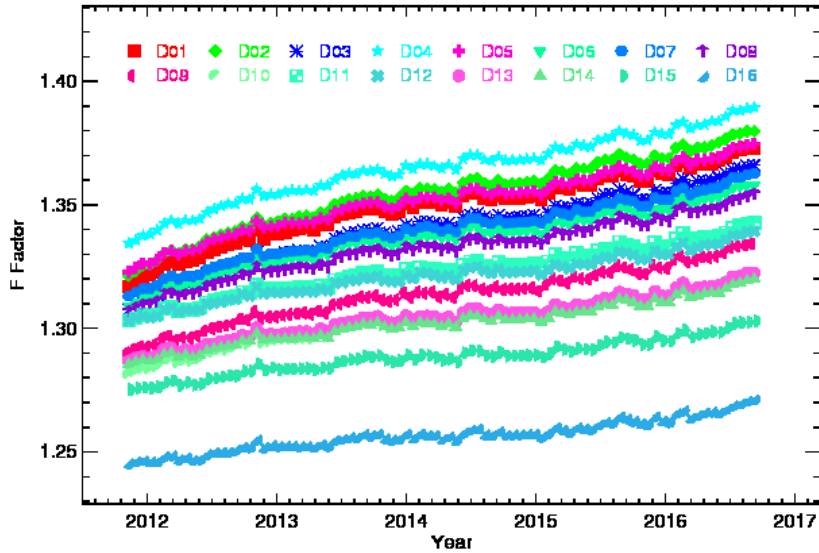
- B, D, M, G : Band, Detector, HAM side, and gain status

J. Sun and M. Wang, “On-orbit calibration of Visible Infrared Imaging Radiometer Suite reflective solar bands and its challenges using a solar,” Appl. Opt., 54, 7210-7223 (2015).

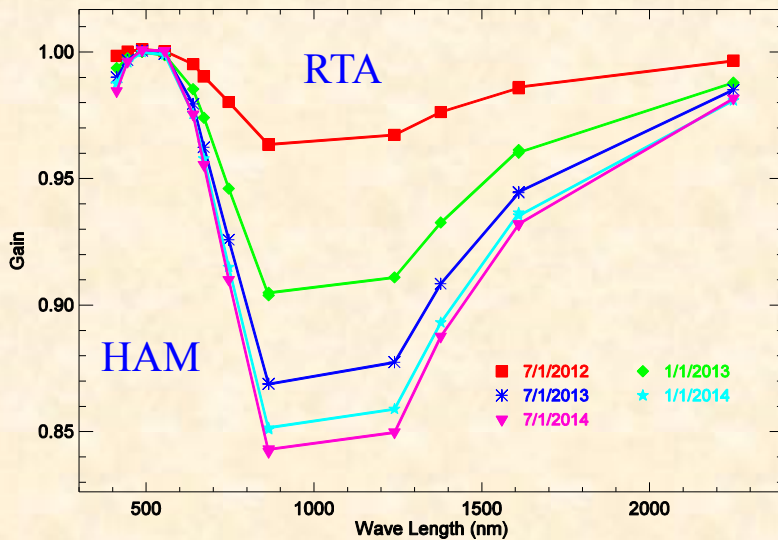
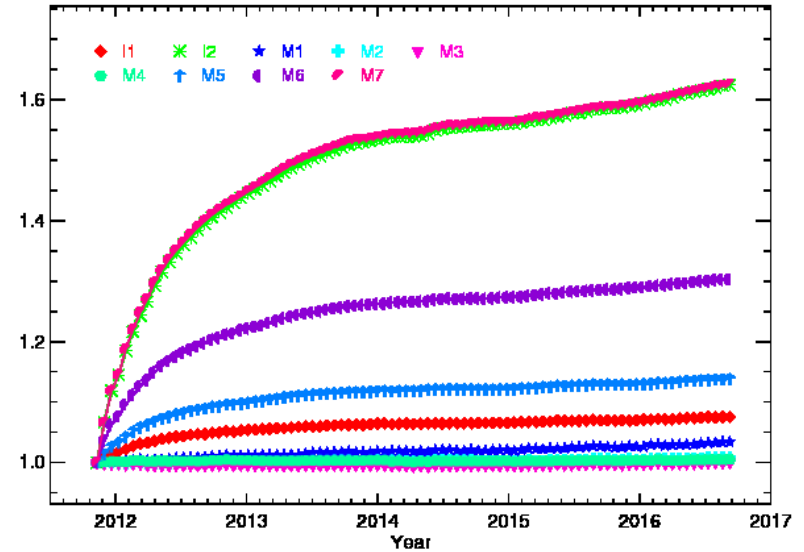


SD Calibration: Every orbit

- **Improvements**
 - **Robust VFs and BRFs from yaw measurements**
 - **Improved H factors**
 - **Sweet spot selection**
 - **Time-dependent RSR**



On-orbit Gain Change



- *Very stable and smooth*
- *Different from MODIS: Much less degradation of the scan mirror*
- *But the input H-factor measured by the SDSM is for the SDSM view direction – KEY ISSUE*

Lunar Calibration Algorithm

- Moon is very stable in its reflectance
- RSB calibration coefficients, F factors, from lunar observations

$$F(B, M) = \frac{g(B)N_{t,M}}{\sum_{D,S,N} L_{pl}(B, D, S, N)\delta(M, M_N)},$$

- $g(B)$: View geometric effect correction (ROLO lunar model and extra correction)

SNPP VIIRS is scheduled to view the Moon approximately monthly (about nine months every year)



- **Advantages**
 - **Lunar surface reflectance has no observable degradation**
 - **Can be used for inter-comparison**

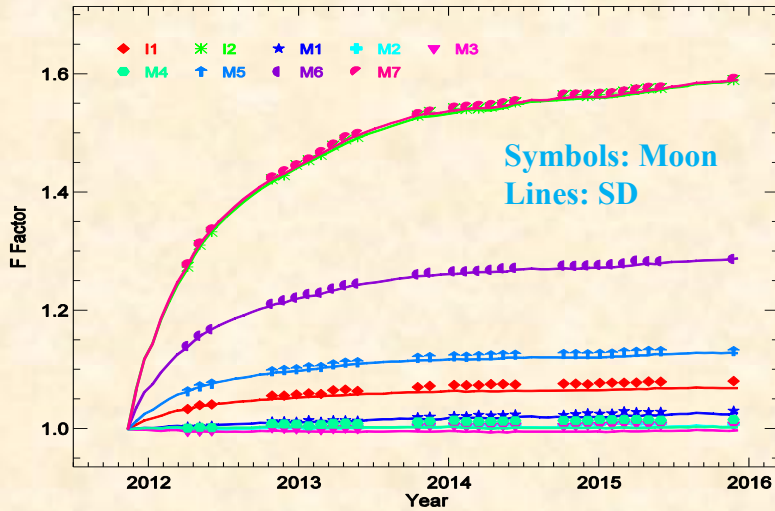
J. Sun, X. Xiong, and J. Butler, “NPP VIIRS on-orbit calibration and characterization using the moon”, Proc. SPIE, 8510,85101I, (2012).

X. Xiong, J. Sun, J. Fulbright, Z. Wang, and J. Butler, “Lunar Calibration and Performance for S-NPP VIIRS reflective Solar Bands”, IEEE Trans. Geosci. Remote Sens., 54, 1052-1061, (2016).

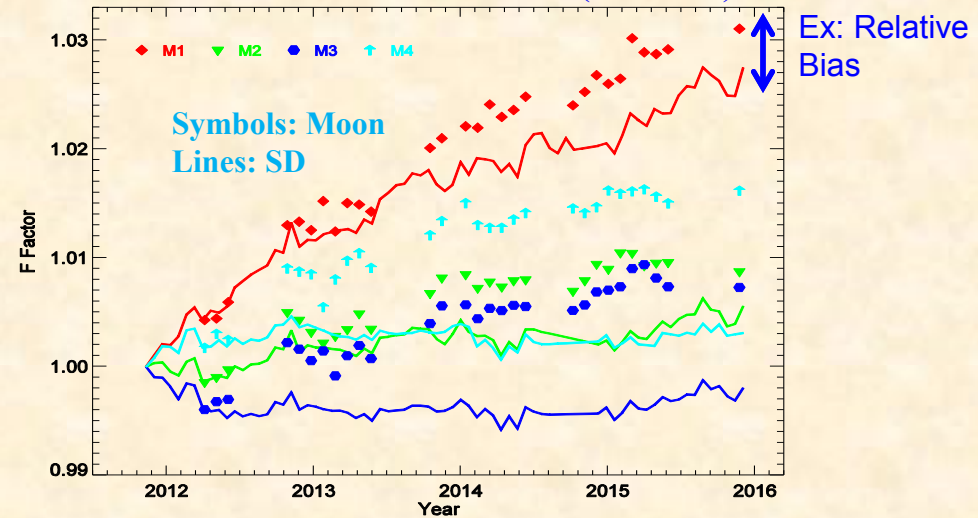
Lunar Calibration Performance

RSB Calibration Coefficients (Lunar F-Factors)

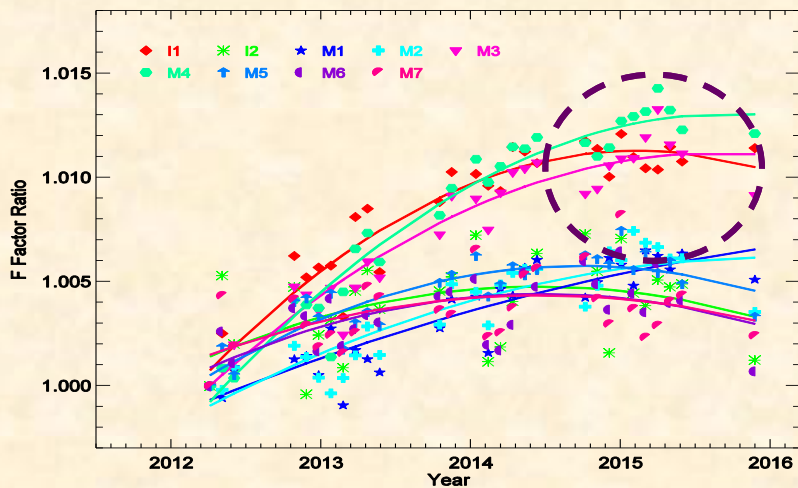
Lunar and SD F Factors



Lunar and SD F factors (M1-M4)

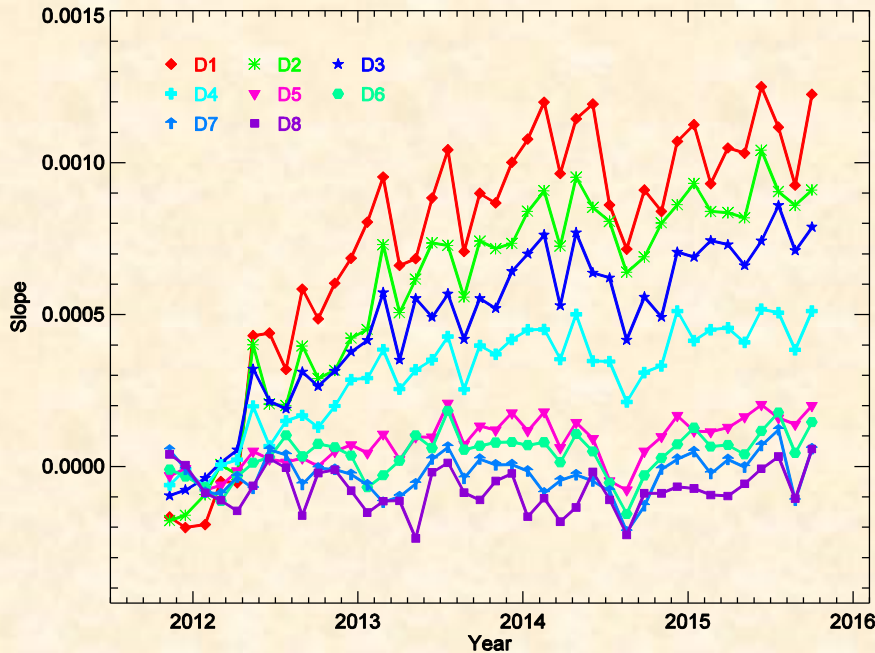


Calibration coefficients Ratios



- *Own Lunar model and correction beyond ROLO model*
- *New Lunar results much improved – smooth, no oscillation - 0.2% stability*
- *SD F-factors and lunar F-factors diverge, especially for short wavelength RSBs.*
- *SD F-factors have error*

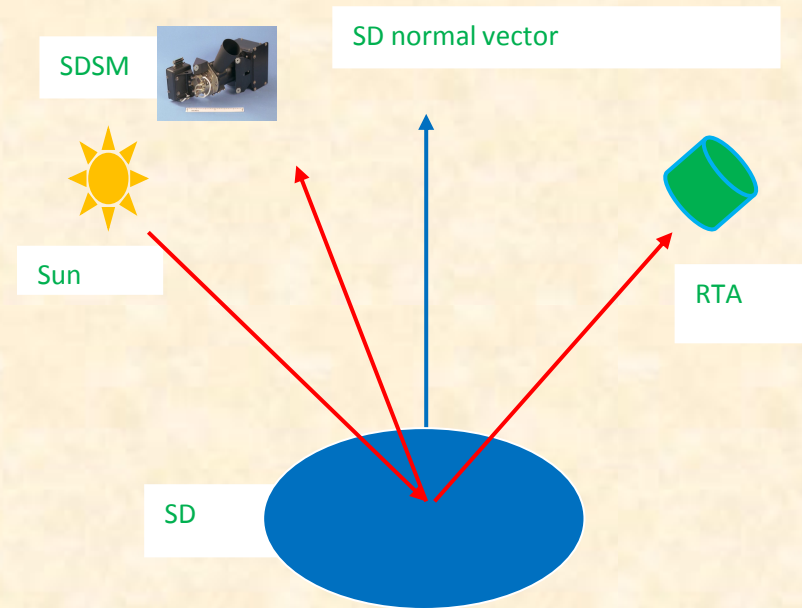
Non-uniformity of SD degradation



Slopes of H-factors in each individual event with respect to solar declination

- It was discovered by 2014 that SD degradation is not uniform
- Standard SD calibration brings non-negligible error into RSB characterization

SDSM and RTA views



J. Sun, M. Chu, M. Wang, "Degradation nonuniformity the solar diffuser bidirectional reflectance distribution factor," Appl. Opt., 55, 6001-6016 (2016).

Hybrid Approach

- SD Calibration
 - SD degrades non-uniformly, resulting long-term drifts
 - Results are stable and smooth
 - Observation in every orbit
- Lunar Calibration
 - No degradation issue
 - Infrequent and no observation in three months every year

- **Hybrid Approach**

$$\mathcal{F}(B, D, M, G) = R(B, t) \cdot F(B, D, M, G)$$

$$R(B, t) = \langle f(B, M, t) \rangle_M / \langle F(B, D, M, 0, t) \rangle_{D, t-15 < t_i < t+15, M}$$

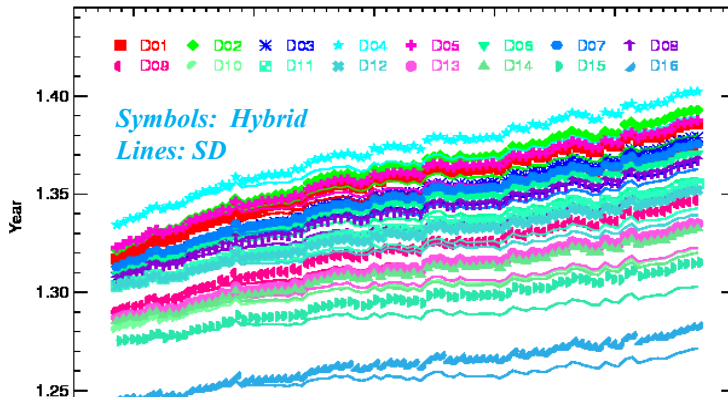
- **Lunar calibration provides long-term baseline**
- **SD calibration provides smoothness and frequency**

F-Factors Ratios are fitted to quadratic polynomials of time

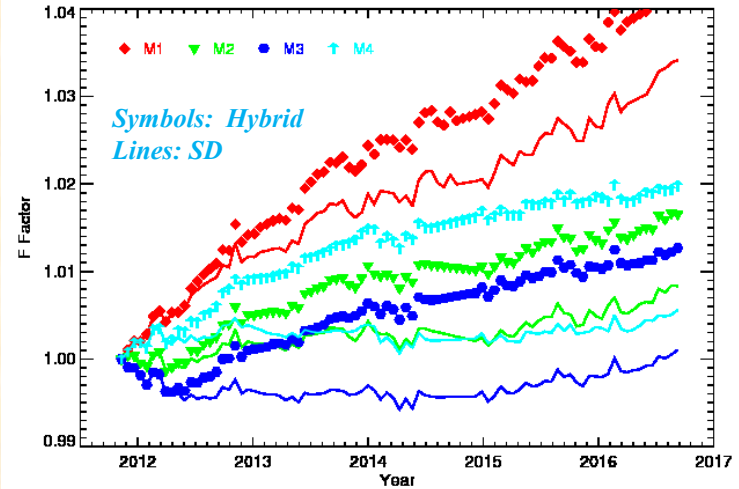
J. Sun and M. Wang, “Radiometric Calibration of the VIIRS Reflective Solar Bands with Robust Characterizations and Hybrid Calibration Coefficients,” Appl. Opt., 54, 9331-9342 (2015).

Hybrid Calibration Performance

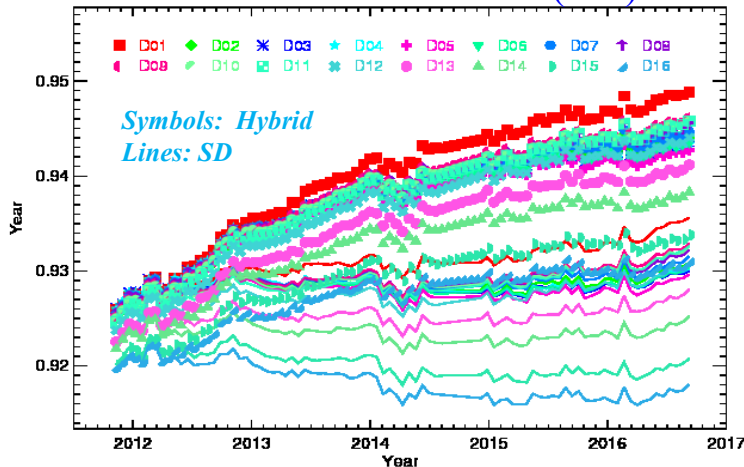
Calibration Coefficients (M1)



Band Averaged (M1-M4)



Calibration Coefficients (M4)



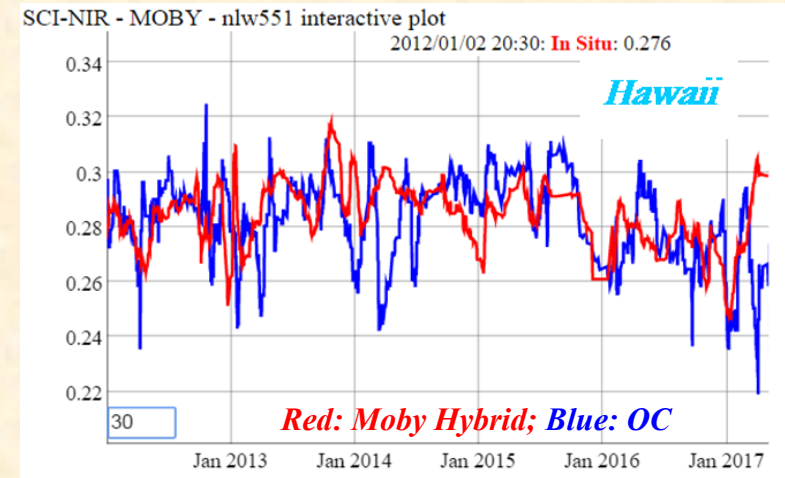
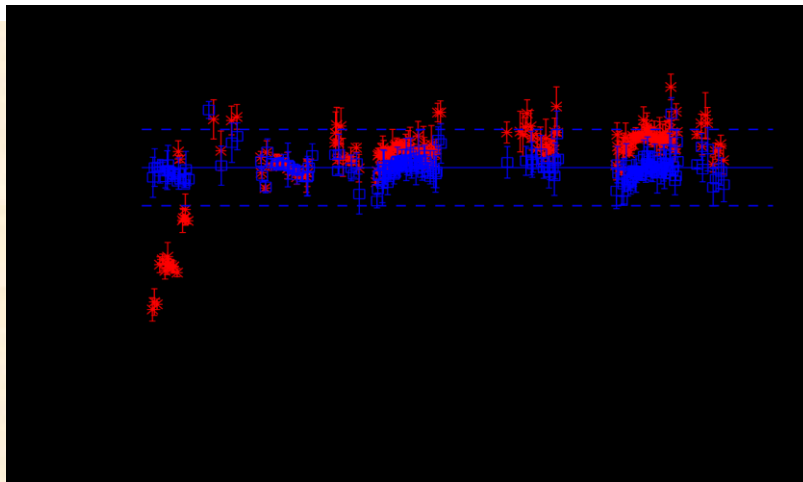
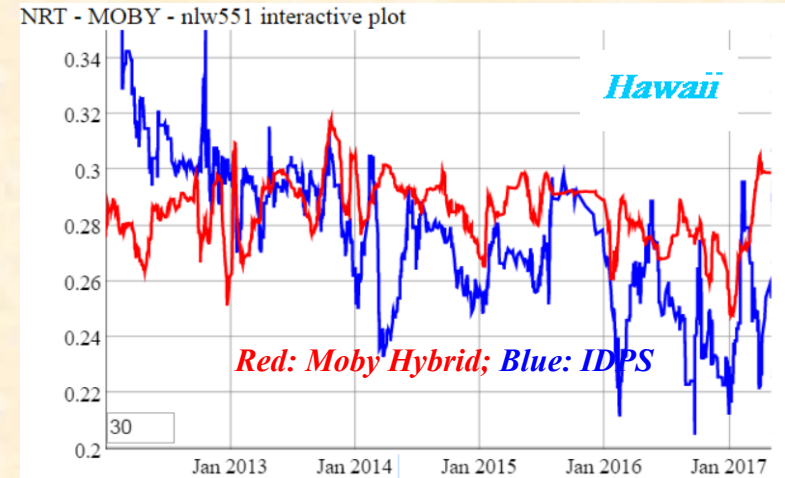
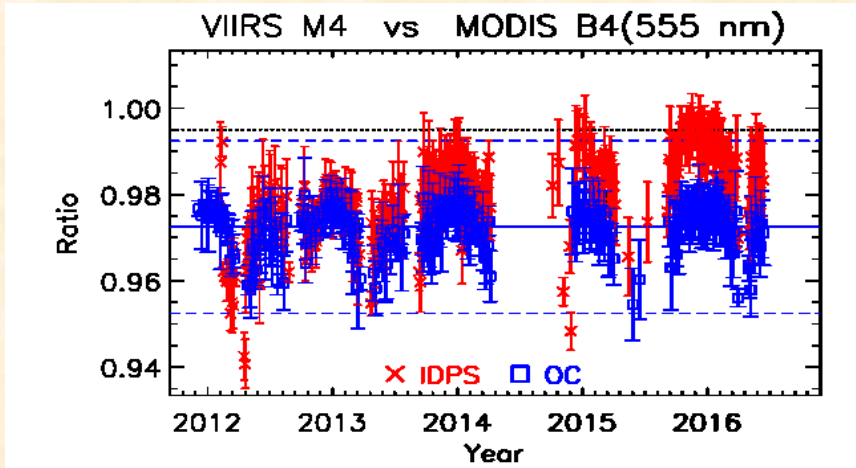
- Hybrid calibration coefficients (Hybrid F-factors) achieves long-term accuracy but also with short-term stability achieving ~0.2% level.
- Earth-based SDR studies show that Hybrid-method indeed mitigated the long-term defect and give stable time-series.

J. Sun and M. Wang, "VIIRS Reflective Solar Bands Calibration Progress and Its Impact on Ocean Color Products," Remote Sensing, 8, 194 (2016).

Inter-sensor and In-situ Comparison

Aqua MODIS and VIIRS Radiance SNO Ratio

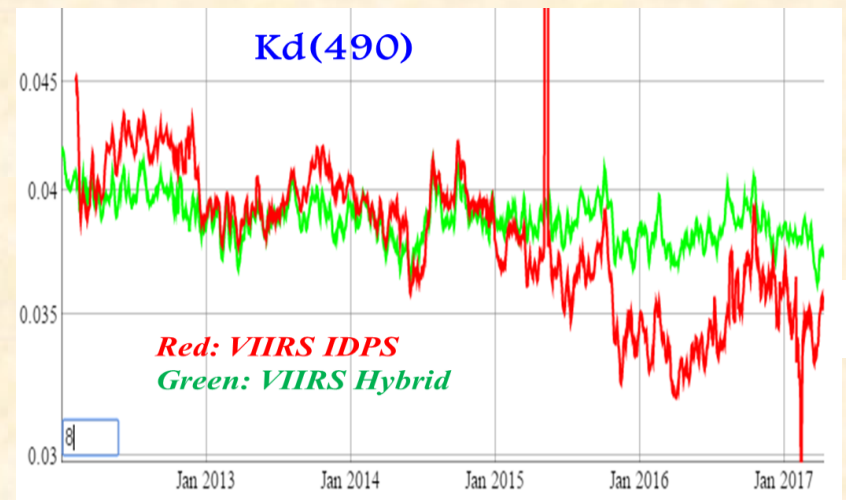
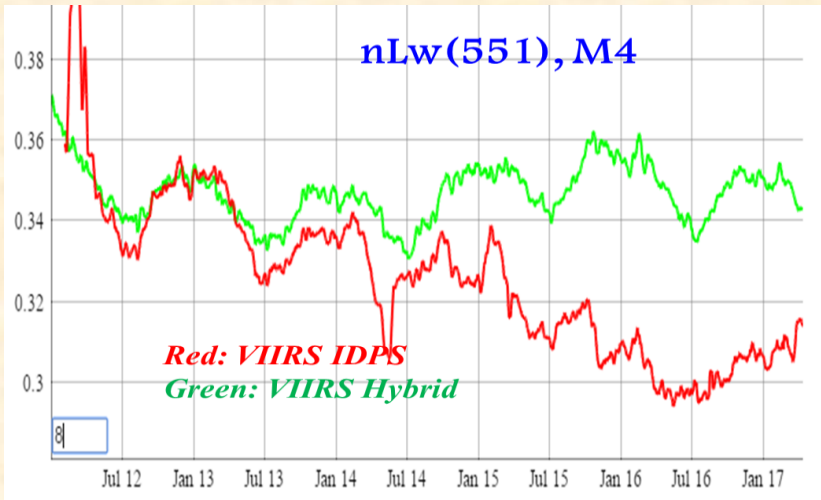
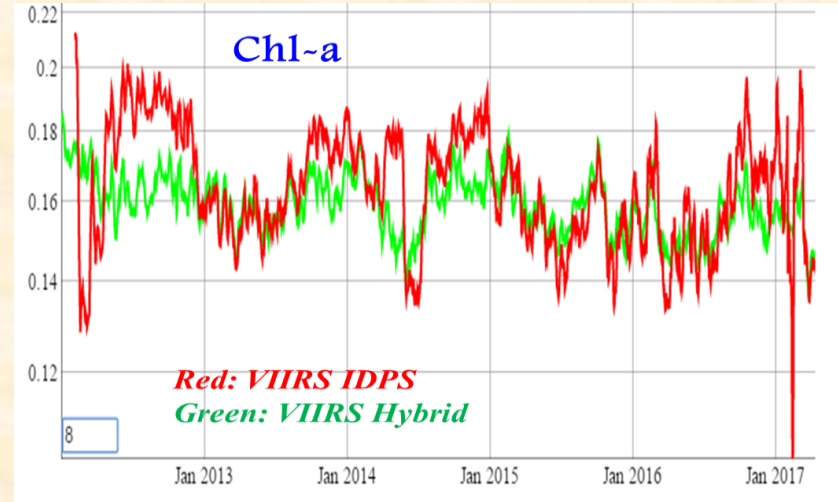
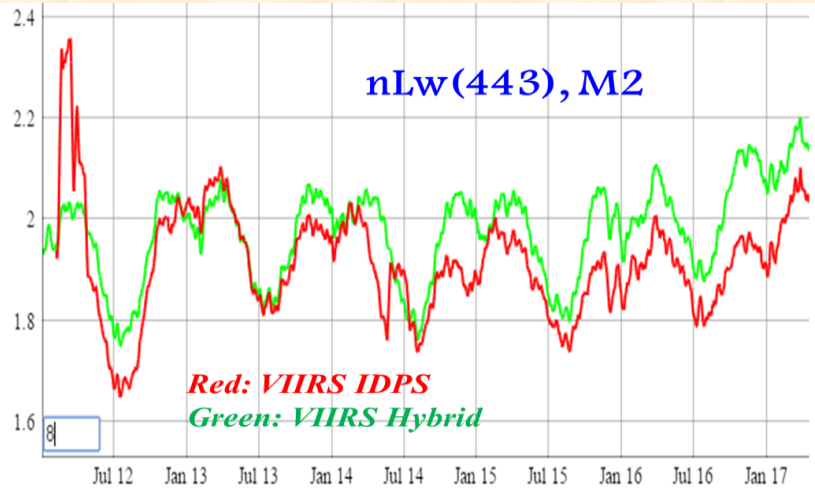
Water Leaving Radiance: nLw(551), M4



Chu, M., J. Sun, and M. Wang, "Radiometric evaluation of the SNPP VIIRS RSB sensor data records via inter-sensor comparison with Aqua MODIS", Proc. SPIE 9972, 99721R (2016).

Ocean Color Products Performance

Global Deep Water (Depth > 1km)



M. Wang, et al, "Evaluation of VIIRS ocean color products," Proc. SPIE 9261, 92610E (2014).

Charts were produced by X. Liu and S. Son.



Summary

- Rigorous calibration of every component is needed. Every component counts.
- RSB calibration to reach 0.2% is not easy but has been done. Results very clean, smooth and **stable**.
- “*SD degradation nonuniformity effect*” brings inherent errors in RSB SD calibration.
- “Hybrid-method” mitigation combining SD and Lunar calibration restores RSB calibration **accuracy**.
- The hybrid coefficients remove long-term bias in ocean color EDR products and enables the VIIRS ocean products for science quality applications. Similar issues is expected in J1-J4 VIIRS.
- VIIRS Ocean Color EDR mission-long data have been reprocessed a couple of times with Hybrid Coef. F-LUTs. Forward delivery of science quality EDR has been implemented since May 2016.
- **Our hybrid calibration coefficients for all RSB were delivered to VIIRS SDR team in September, 2016 for operational VIIRS SDR reprocessing.**
- *We anticipate more challenging issues to come and we are preparing.*



Backup

Table 1. Specification for SNPP VIIRS RSBs and SDSM detectors.

VIIRS Band	CW* (nm)	Band Gain	Detectors	Resolution*	SDSD Detector	CW* (nm)
M1	410	DG	16	742m x 776m	D1	412
M2	443	DG	16	742m x 776m	D2	450
M3	486	DG	16	742m x 776m	D3	488
M4	551	DG	16	742m x 776m	D4	555
I1	640	SG	32	371m x 387m	NA	NA
M5	671	DG	16	742m x 776m	D5	672
M6	745	SG	16	742m x 776m	D6	746
M7	862	DG	16	742m x 776m	D7	865
I2	862	SG	32	371m x 387m	D7	865
NA	NA	N	16		D8	935
M8	1238	SG	16	742m x 776m	NA	NA
M9	1378	SG	16	742m x 776m	NA	NA
M10	1610	SG	16	742m x 776m	NA	NA
I3	1610	SG	32	371m x 387m	NA	NA
M11	2250	SG	16	742m x 776m	NA	NA

*CW: Center Wavelength; DG: Dual Gain; SG: Singla Gain; Resolution: Track x Scan at Nadir after aggregation