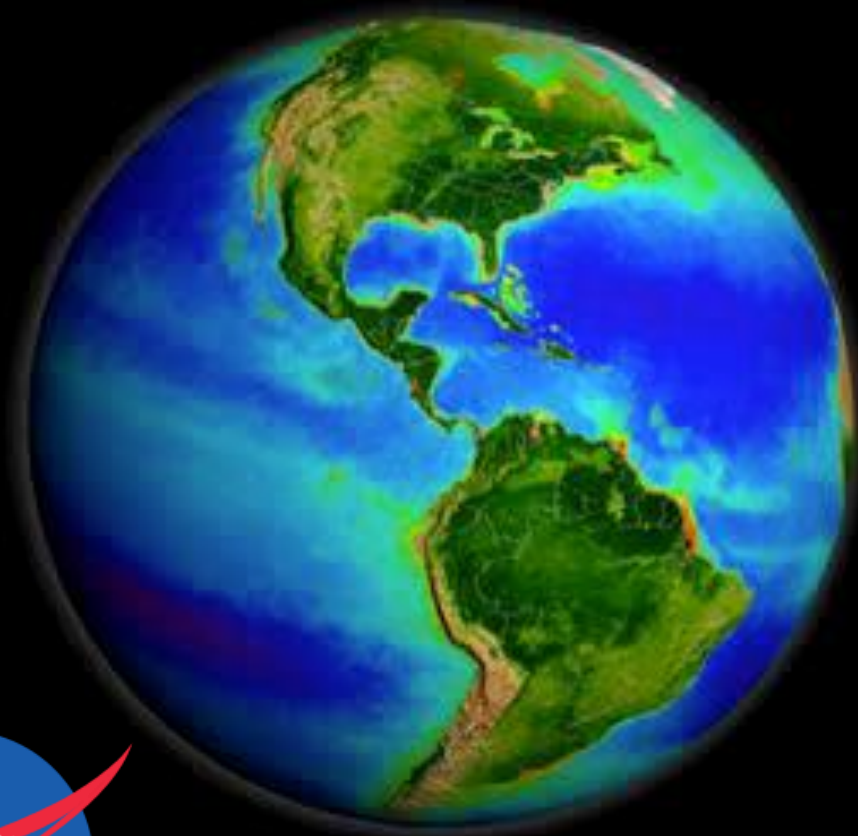


Atmospheric correction of hyperspectral ocean color sensors: application to HICO



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Remote sensing of the Atmosphere-Ocean system

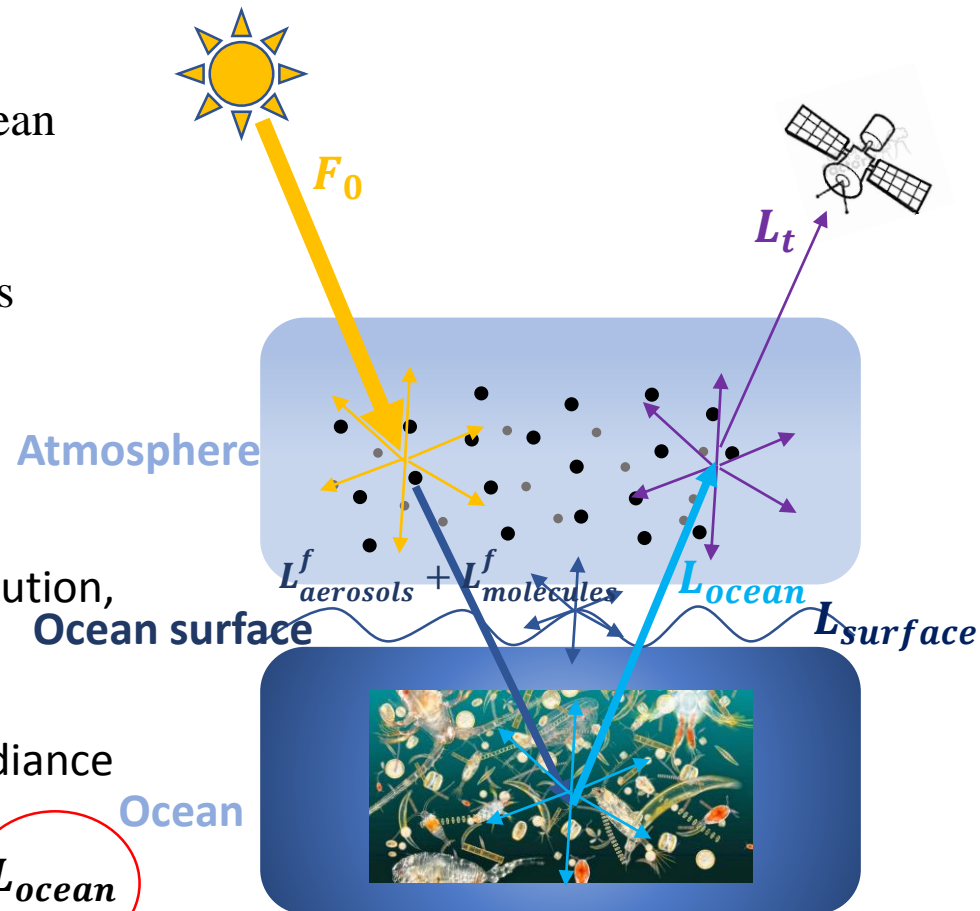
Solar Light interacts with the Atmosphere-Ocean system according to the following principles:

- Scattering by Aerosols, hydrosols and gases
- Absorption by Aerosols, hydrosols and gases
- Reflection and refraction of the air-sea interface

For simplicity, we assume the radiance contribution, L , of every component of the AO system to be additive:

➤ Satellite observations measure the total radiance at the top of atmosphere, L_t

➤ $L_t = L_{aerosols} + L_{molecules} + L_{surface} + L_{ocean}$

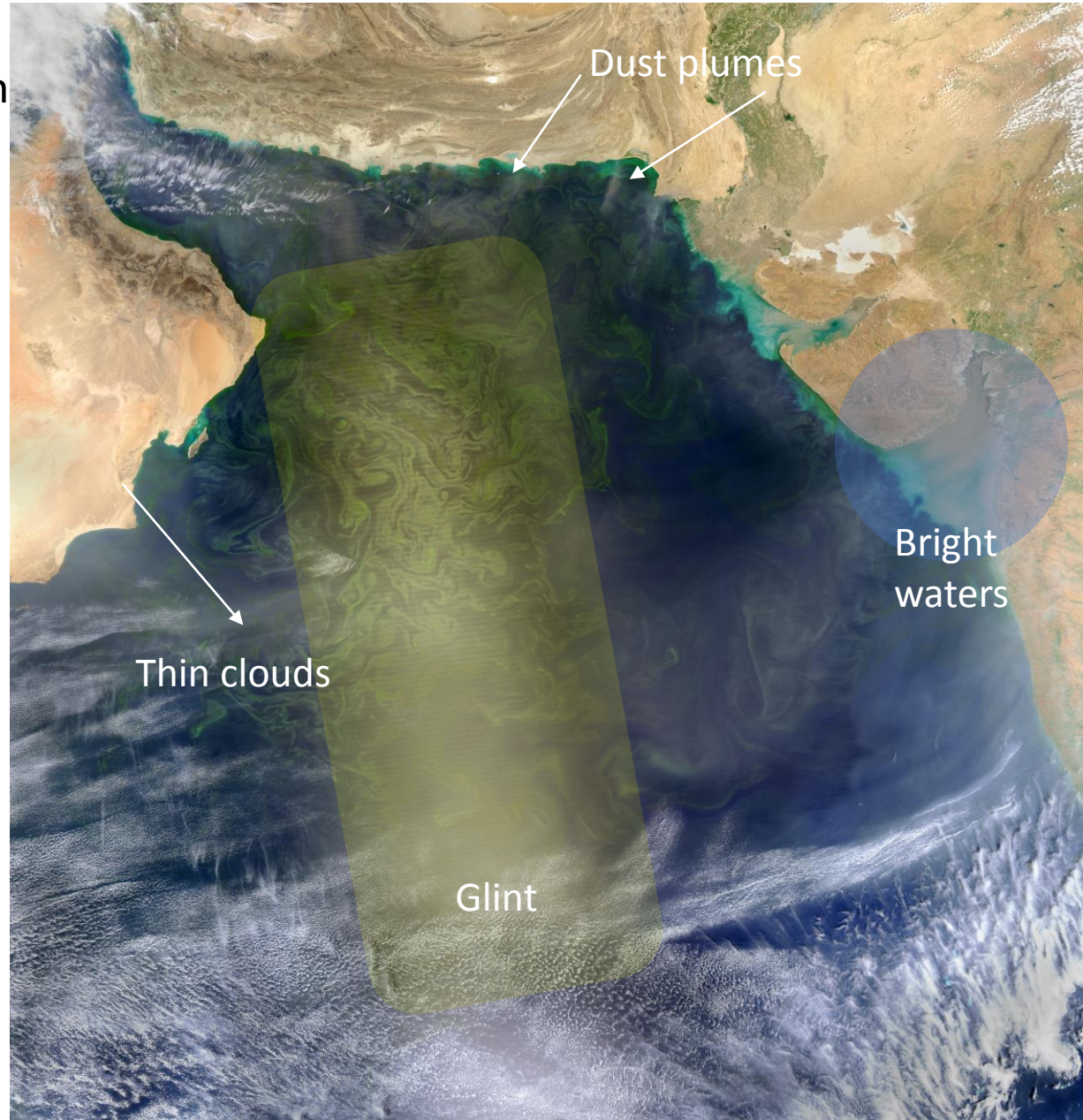


Ocean color enthusiasts need to measure this signal. Is that simple? NO



MODIS Aqua scene of the Arabian sea

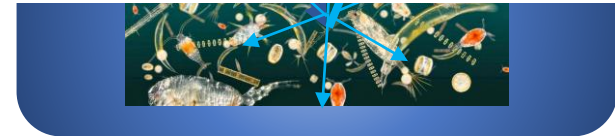
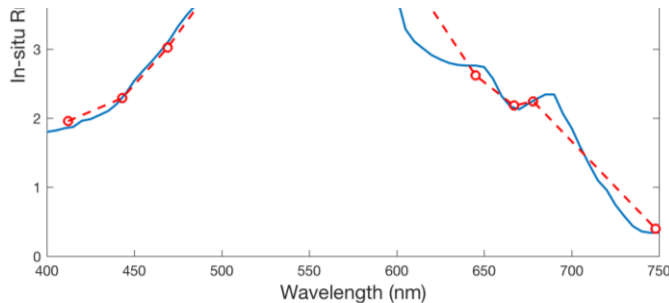
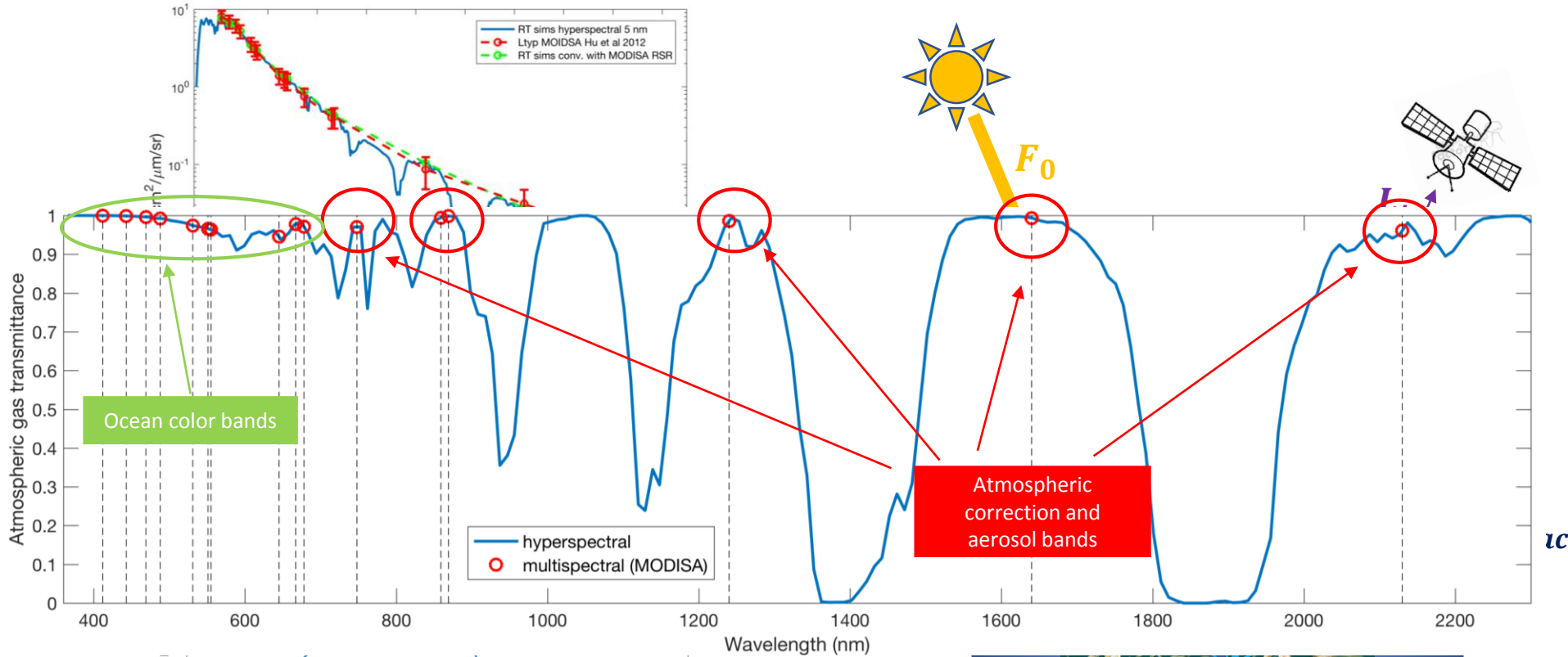
- Natural scenes acquired from space can be complex:
 - Absorbing aerosols
 - Strong Sun glint
 - Bright waters
 - Thin clouds
- Several atmospheric correction techniques are used in NASA's ocean color operational algorithm to mitigate these problems



Heritage Multi-spectral spaceborne sensors

MODIS Aqua example:

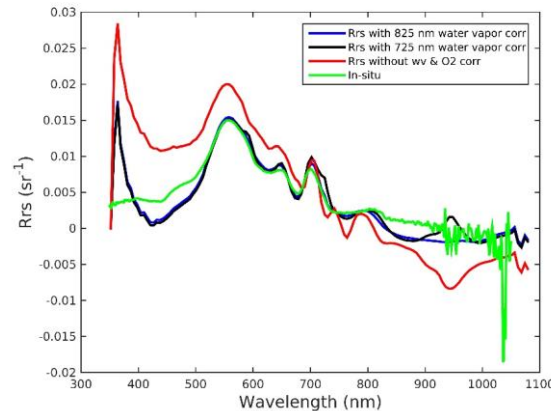
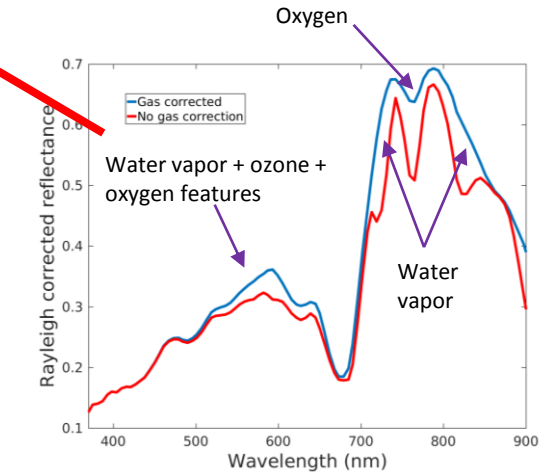
- Spectral information content:



Ocean color, AC and aerosols bands are located at Atmospheric windows (i.e., minimally impacted by absorbing gases)

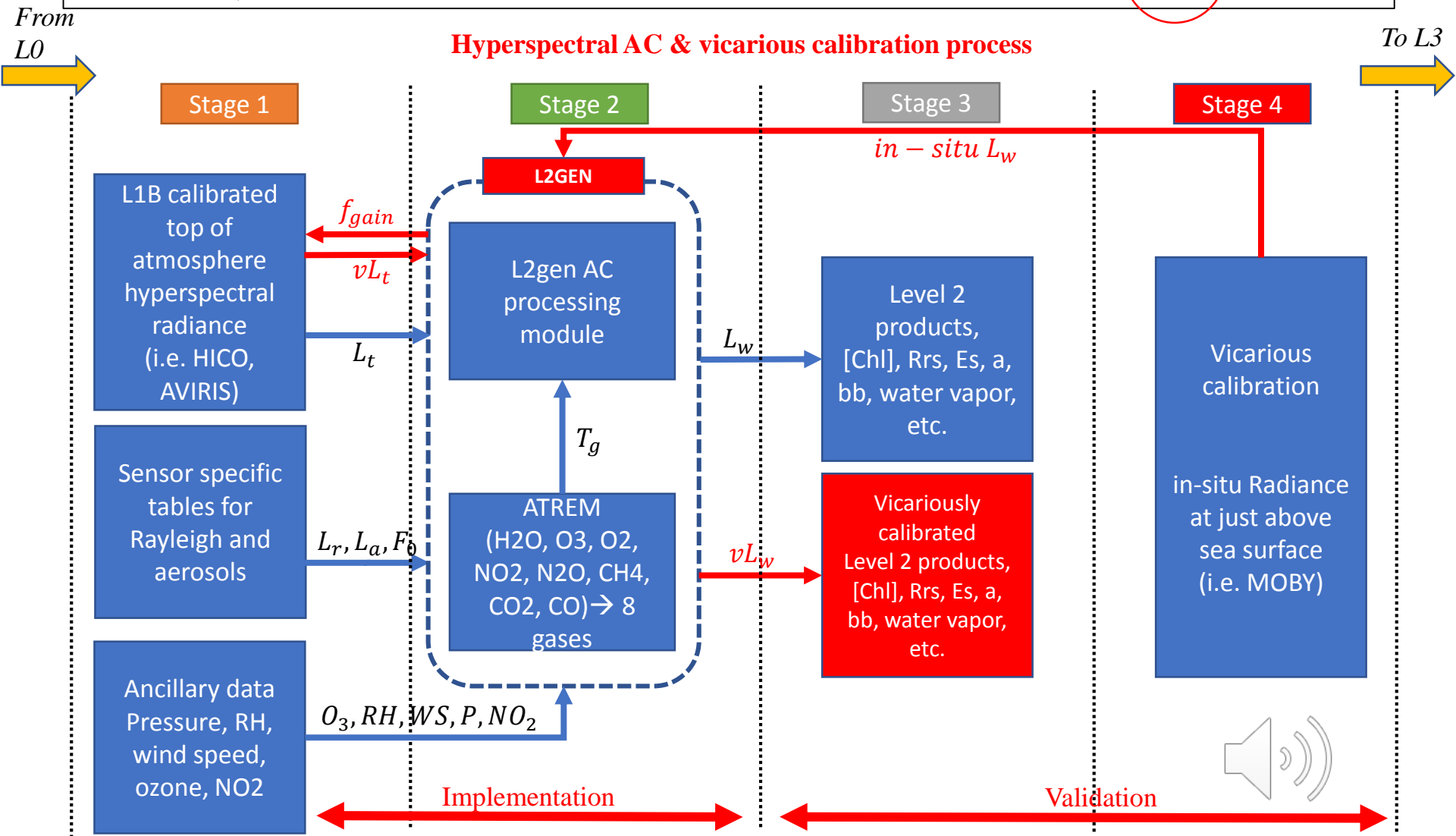
Why do we need to compensate for absorbing gases in the AC?

- Absorbing gases including: water vapor, oxygen, ozone and nitrogen dioxide modulate the measured TOA radiance significantly within the visible spectrum.
- A correction algorithm for gases is needed to remove the unwanted spectral features in ocean reflectance.
- Erroneous correction of gases can significantly degrade ocean color data quality and plankton type algorithms.

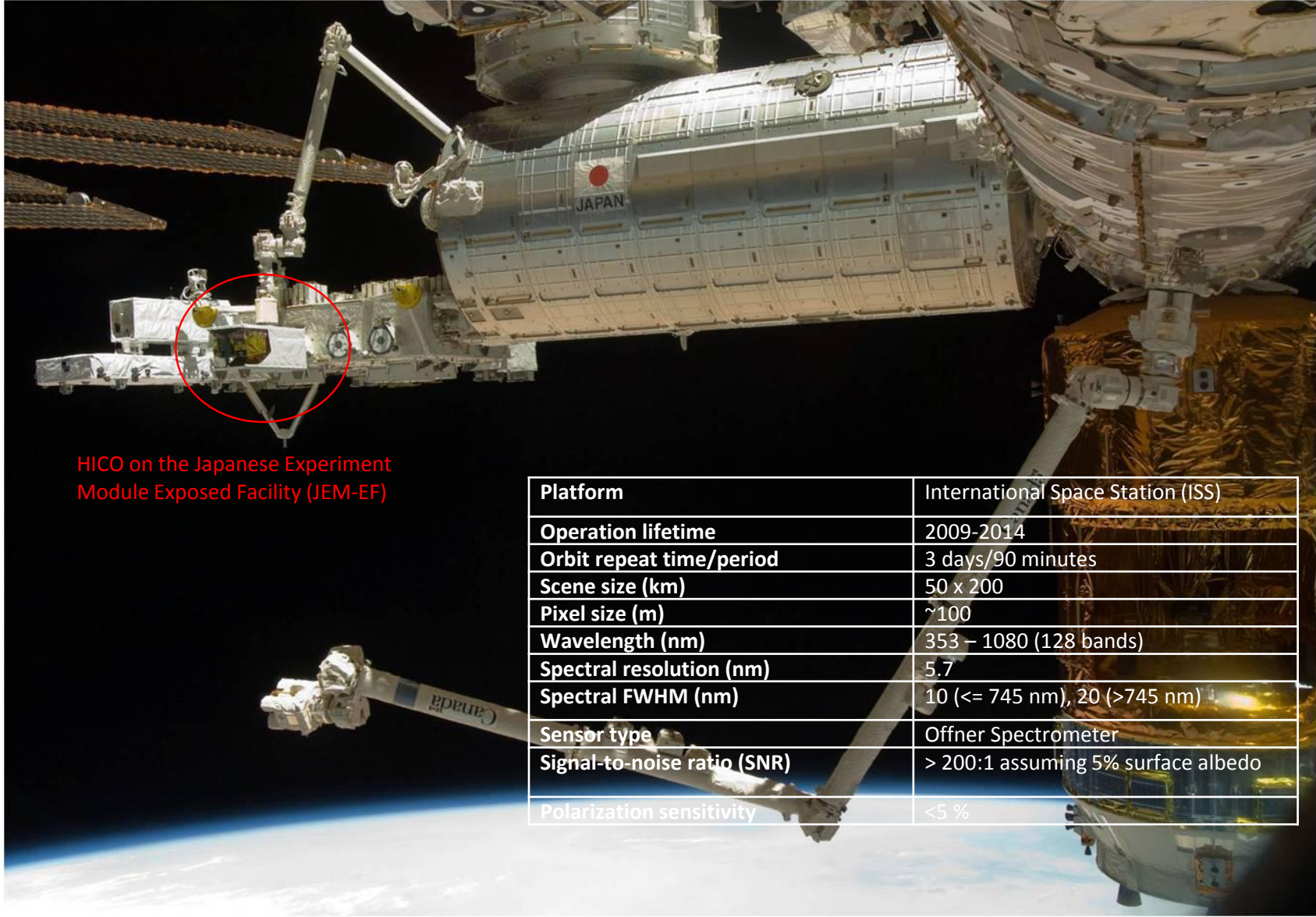


NASA's operational multispectral AC algorithm extended to hyperspectral

$$L_t(\lambda) = (L_r(\lambda) + L_a(\lambda) + L_{ra}(\lambda) + t(\lambda)L_f(\lambda) + T(\lambda)L_g(\lambda) + t(\lambda)L_w(\lambda)) \times T_g(\lambda)$$



Hyperspectral Imager for Coastal Ocean (HICO)



HICO on the Japanese Experiment Module Exposed Facility (JEM-EF)

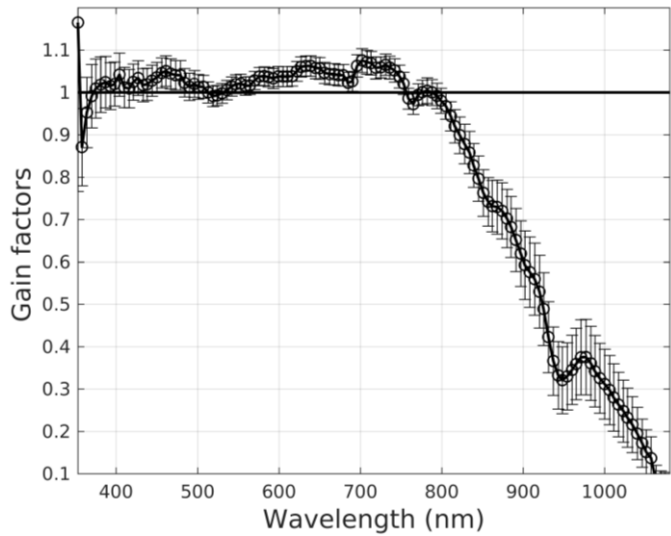
| | |
|------------------------------------|--------------------------------------|
| Platform | International Space Station (ISS) |
| Operation lifetime | 2009-2014 |
| Orbit repeat time/period | 3 days/90 minutes |
| Scene size (km) | 50 x 200 |
| Pixel size (m) | ~100 |
| Wavelength (nm) | 353 – 1080 (128 bands) |
| Spectral resolution (nm) | 5.7 |
| Spectral FWHM (nm) | 10 (≤ 745 nm), 20 (>745 nm) |
| Sensor type | Offner Spectrometer |
| Signal-to-noise ratio (SNR) | $> 200:1$ assuming 5% surface albedo |
| Polarization sensitivity | $< 5\%$ |



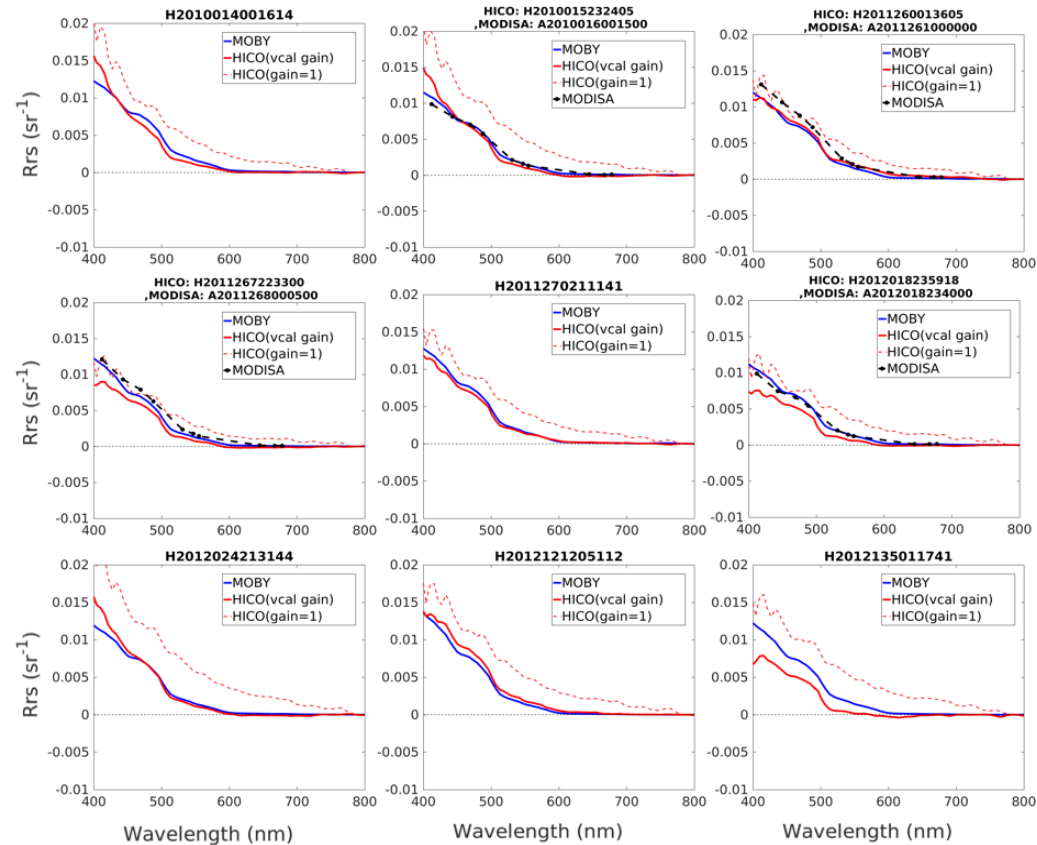
Vicarious calibration of HICO

- Because of HICO's calibration problems, such as thermal instability, second order diffraction effects, and lack of an on-board calibration system, a vicarious calibration of HICO is needed to improve ocean color data quality

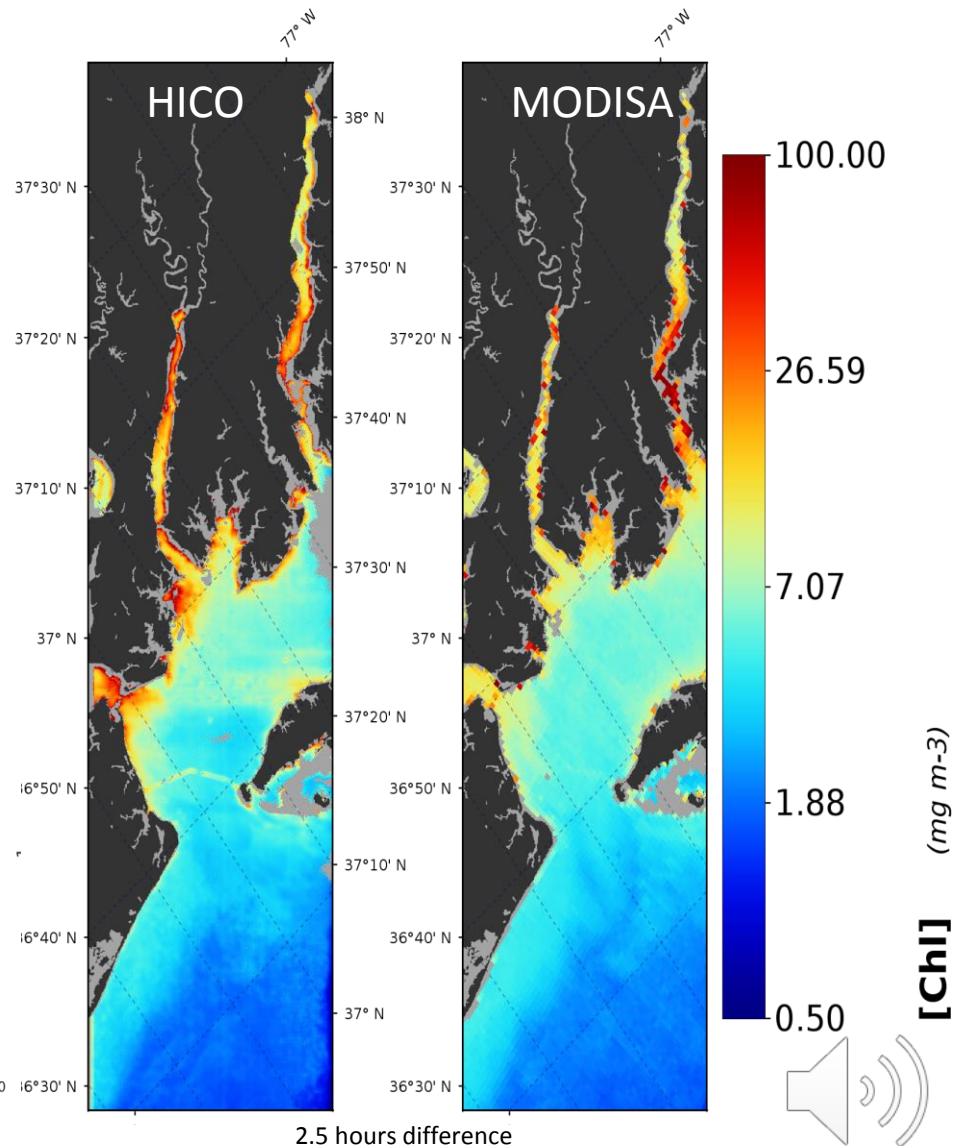
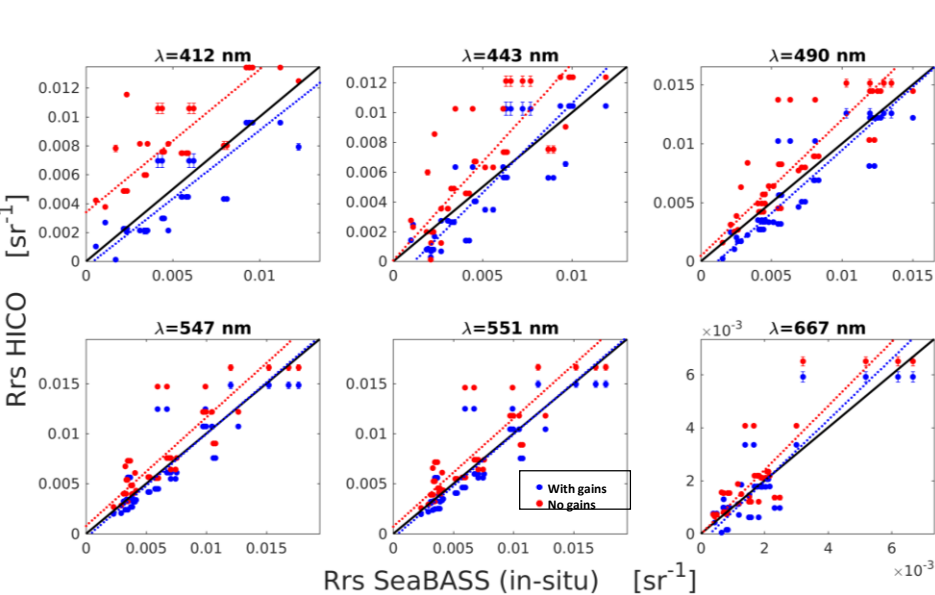
Based on co-incident, co-located in-situ MOBY data, vicarious gains were derived



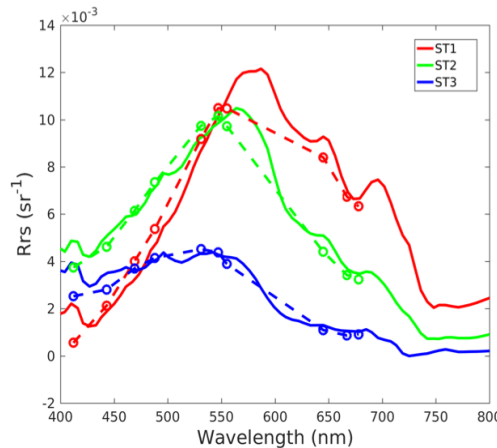
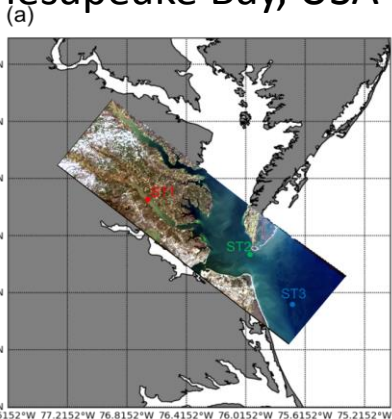
At MOBY



Global in-situ to HICO validation and MODISA to HICO comparison



Chesapeake Bay, USA



Conclusion

- Hyperspectral ocean color remote sensing requires a proper compensation for absorbing gases in the atmosphere.
- NASA's operational multispectral AC algorithm has been extended to hyperspectral to include compensation of all gases, especially water vapor and oxygen.
- After the vicarious calibration of HICO, derived ocean color products were improved.
- The hyperspectral AC algorithm for HICO is currently available through SeaDAS version 7.4 on <https://oceancolor.gsfc.nasa.gov>



Thank you for listening!